

Two-Photon Interferometry for Wide-Angle Precision Astrometry

Aaron Muenninghoff - PhD student

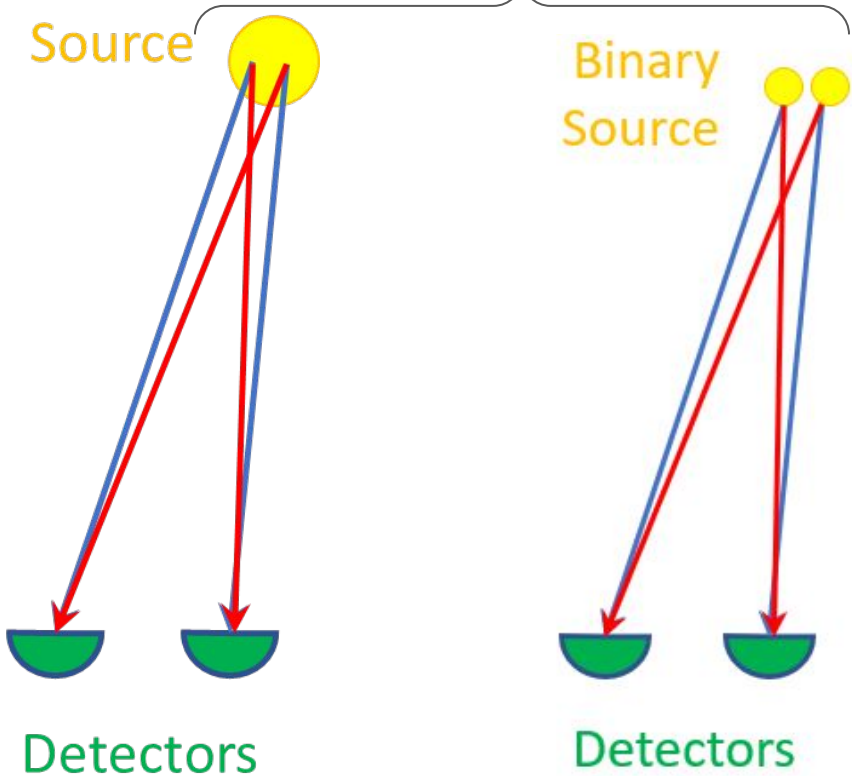
Stony Brook University & **Brookhaven National Lab**

Future Prospects of Intensity Interferometry - Perimeter Institute

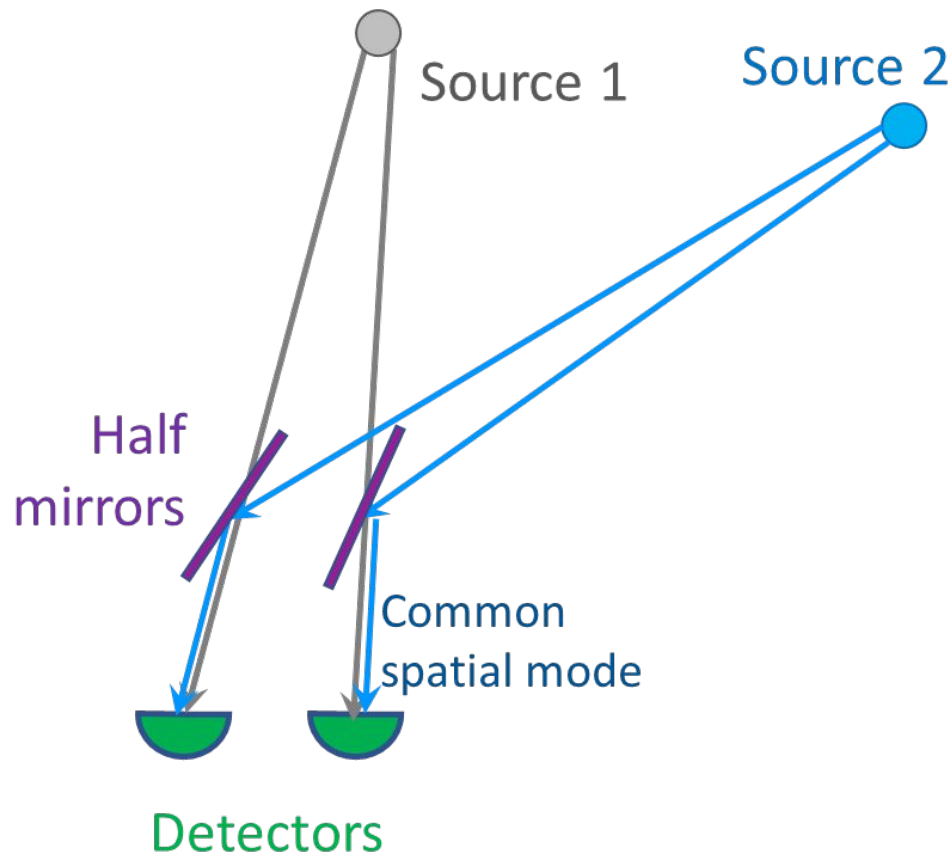
10/30/24

HBT Generalization

Standard HBT



Two-Photon Interferometry for Wide-angle Precision Astrometry



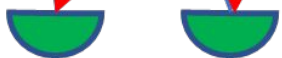
HBT Generalization

Standard HBT

Source



Binary Source

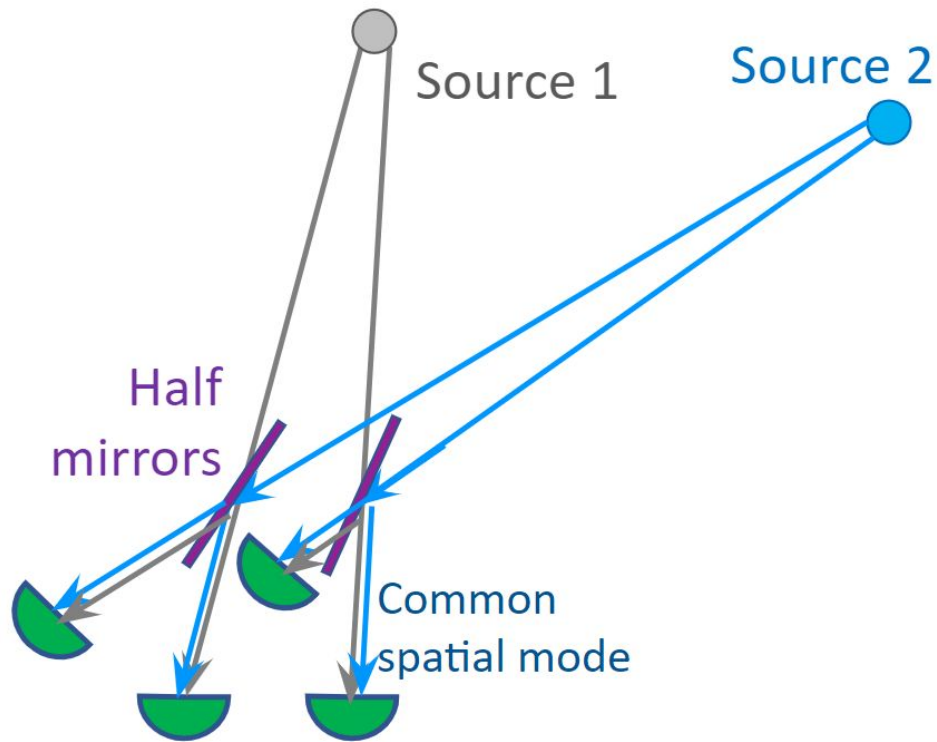


Detectors



Detectors

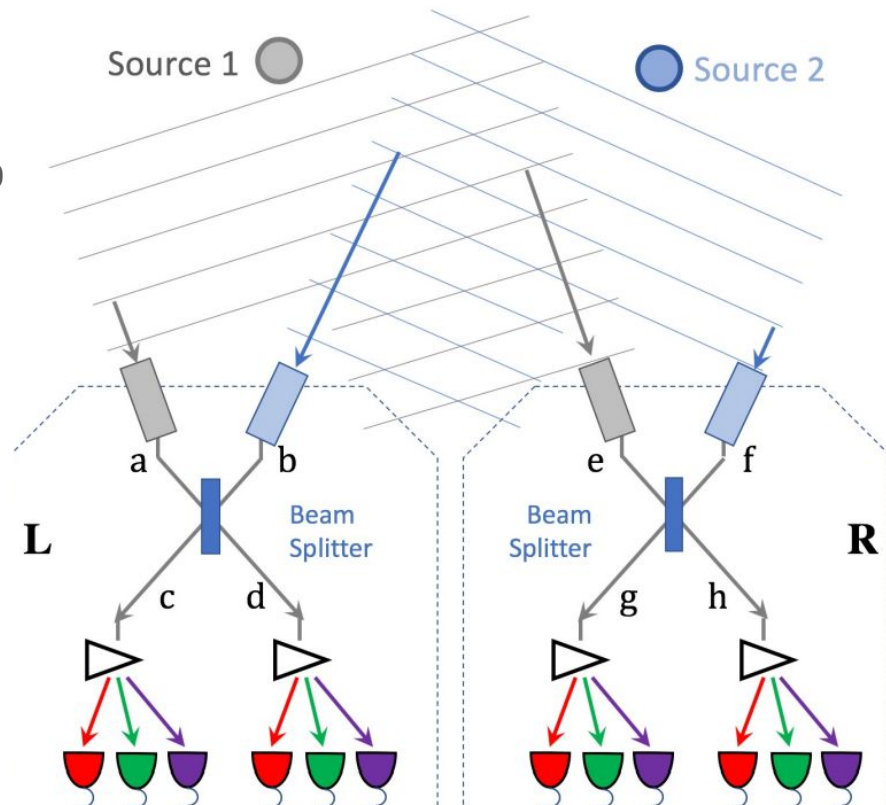
Two-Photon Interferometry for Wide-angle Precision Astrometry



Detectors

The SNSV Scheme

- Original paper: Stankus et al. (2022) *Open Journal of Astrophysics* doi:10.21105/astro.2010.09100
- Two stations with no optical connection
 - Longer baselines, higher precision
- Astrometric measurement (dynamic)
 - Parallax distance measurement
 - Exoplanet-induced stellar reflex motion
 - Stellar spin axis (photocenter movement)
 - μHz gravitational wave detection

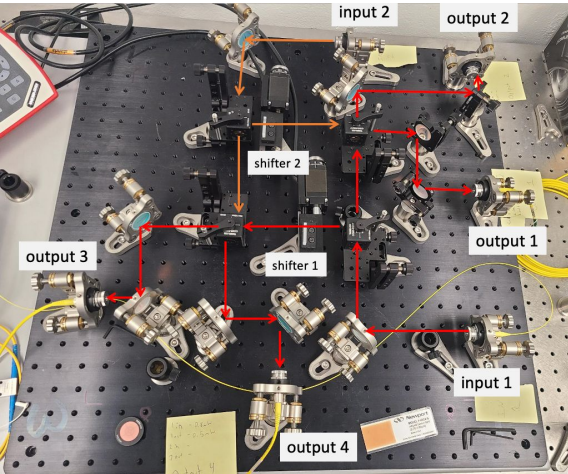


Stankus et al. 2022

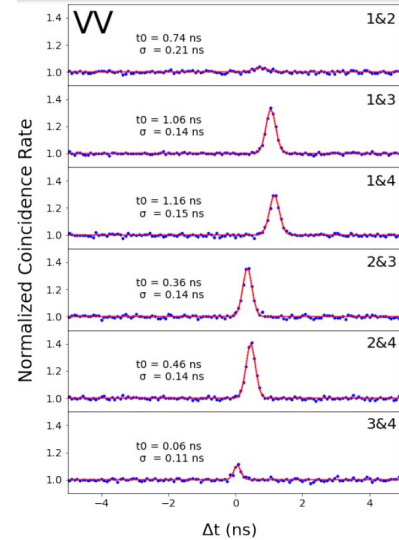
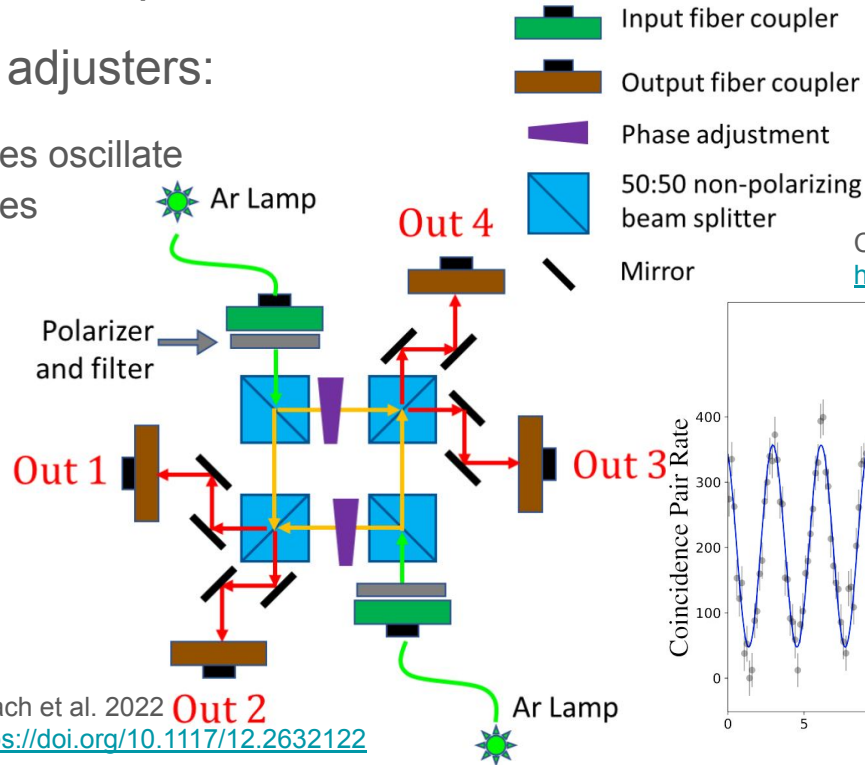
<https://doi.org/10.21105/astro.2010.09100>

SNSV Scheme: Tabletop demonstration

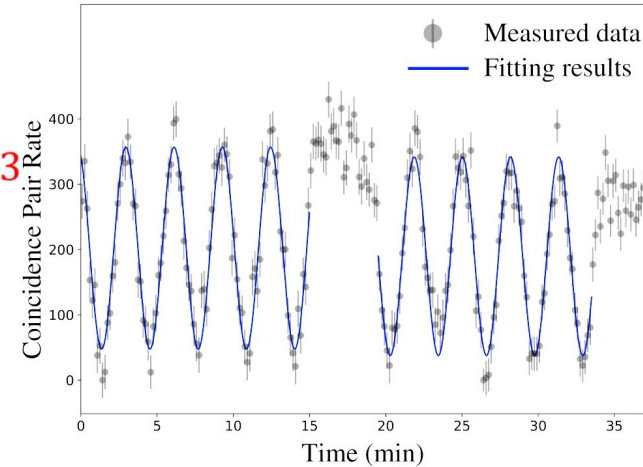
- Tested principle in a tabletop demonstration
- When moving phase adjusters:
 - Cross-station pair rates oscillate
 - Same-station pair rates remain constant



Keach et al. 2022 <https://doi.org/10.1117/12.2632122>



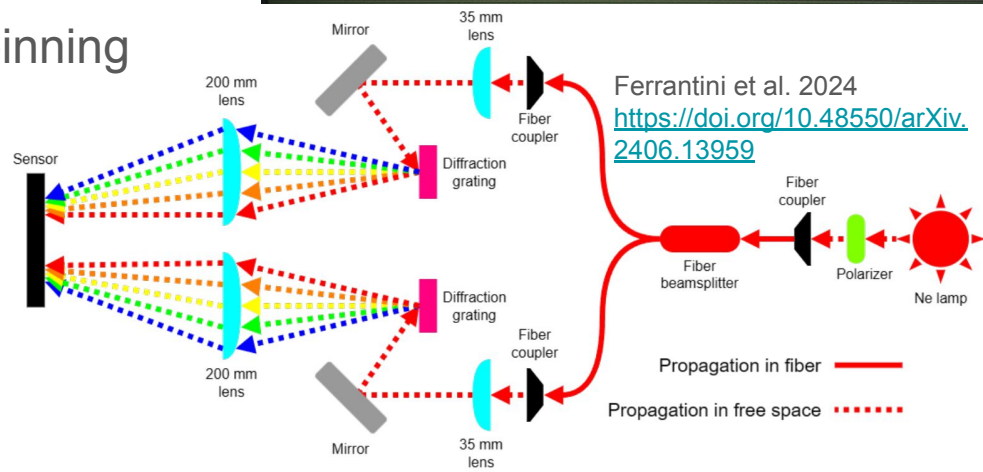
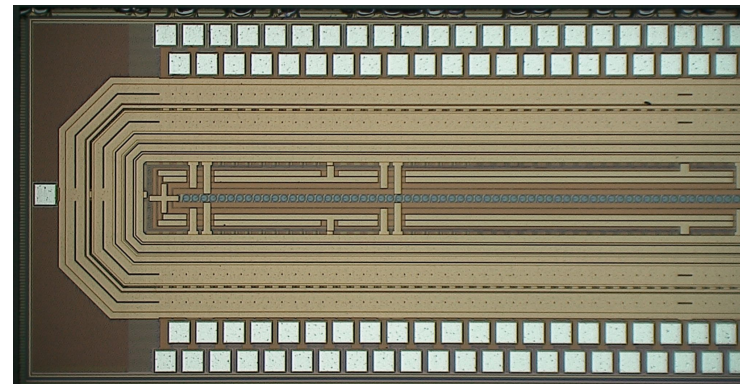
Crawford et al. 2023
<https://doi.org/10.1364/OE.486342>



LinoSPAD2 Fast Spectrometer

Milanese et al. 2023
<https://doi.org/10.1364/OE.505748>

- LinoSPAD2: 1x512 linear array of Single-Photon Avalanche Diodes (SPAD)
 - Made by EPFL in Switzerland
 - ~50ps RMS timing resolution (single photon)
 - Low dark count rate
 - Room temperature operation
- Dual spectrometer key to spectral binning
 - 0.1nm/pix spectral scale
- $(\Delta t)(\Delta E) \sim 10 * (\hbar/2)$
 - Benchmark, not direct comparison

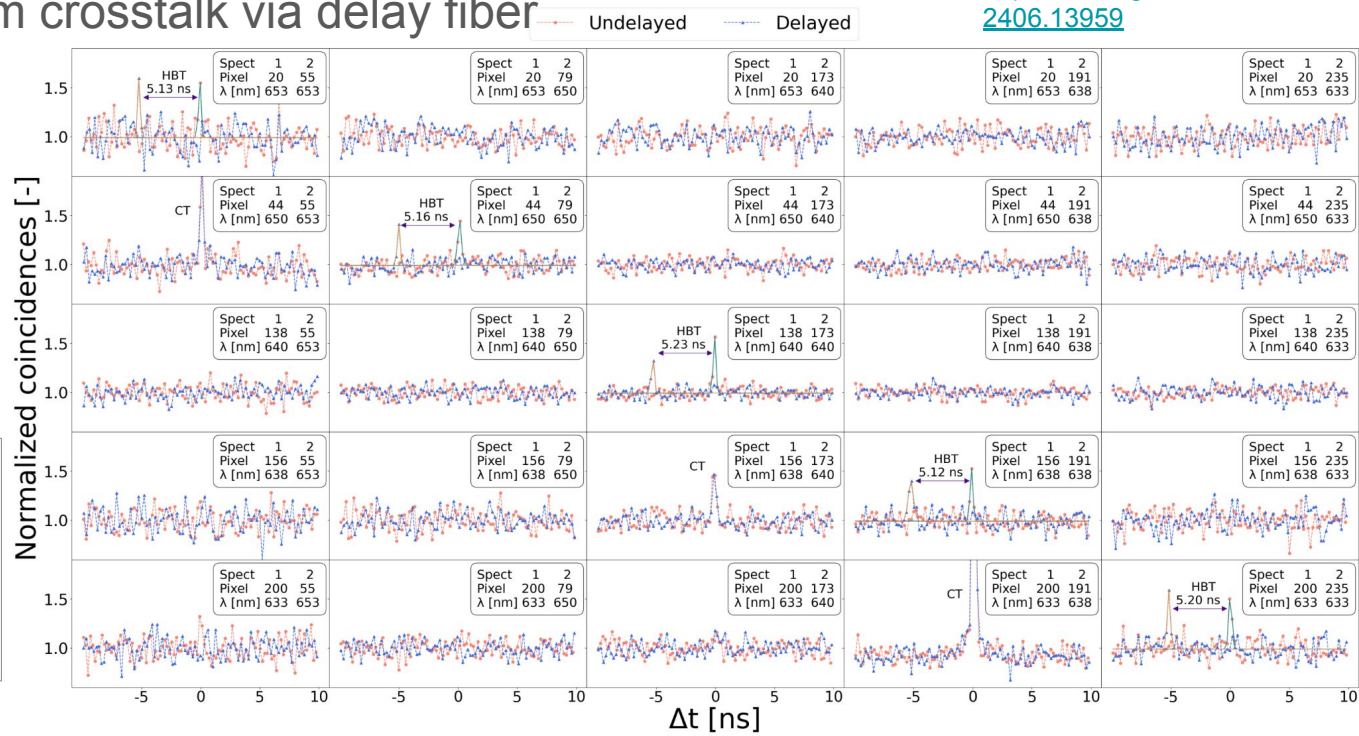
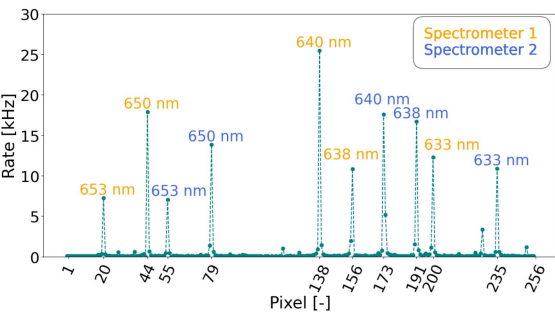


LinoSPAD2 Fast Spec.: Multifrequency HBT

- Simultaneous detection of HBT for 5 Neon lines
- Disambiguate from crosstalk via delay fiber
- Working towards broadband HBT measurement

Ferrantini et al. 2024

<https://doi.org/10.48550/arXiv.2406.13959>

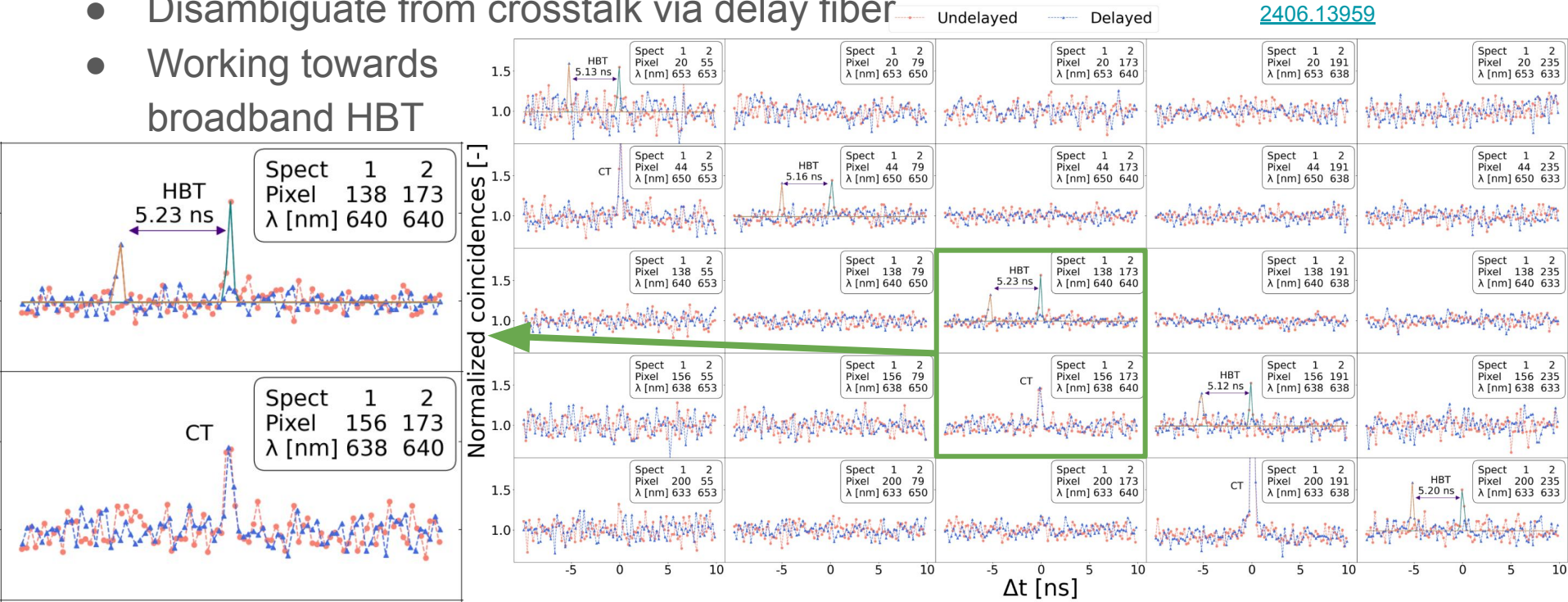


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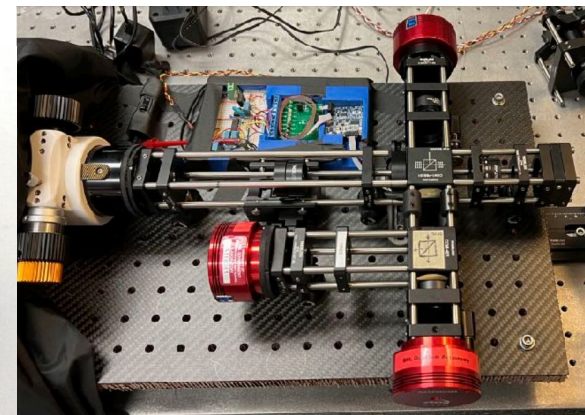
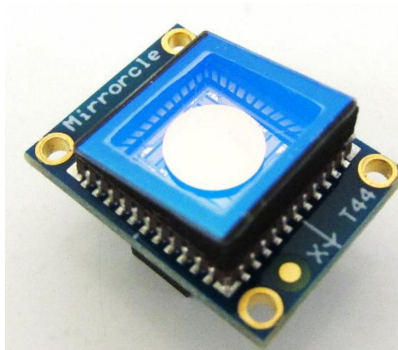
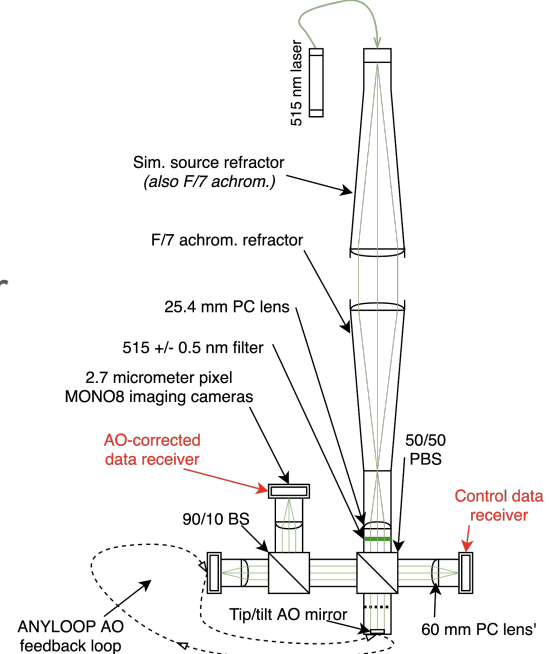
Ferrantini et al. 2024

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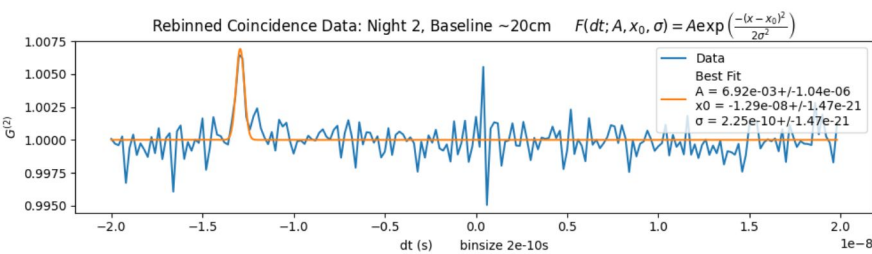
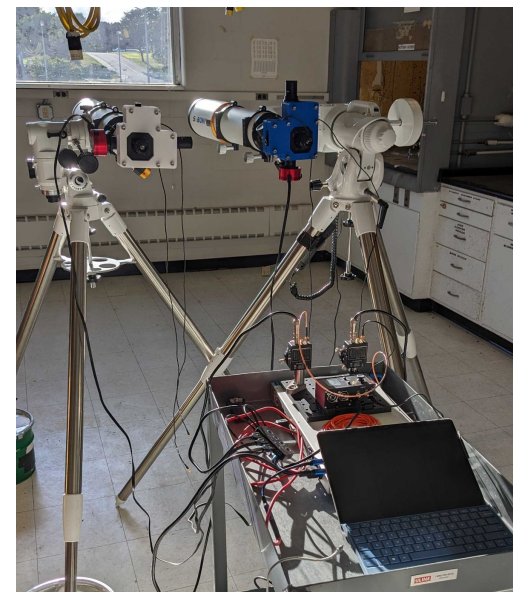
Collecting Single-mode Starlight

- Require light in same spatial mode => single-mode fiber
- First-order adaptive optics correct for atmospheric PSF position movement
- Mirrorcle fast steering mirror
 - Controlled with Anyloop, an open-source control loop software (<https://github.com/cdqp/anyloop>)



Ground Tests with Artificial LED “Star”

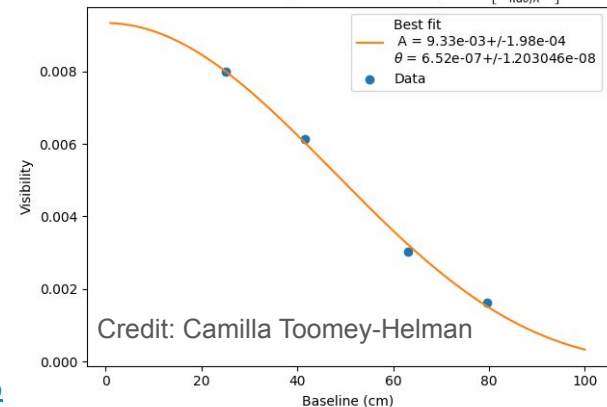
- “Artificial star”: D=50μm fiber emitting filtered & polarized light from a red LED
 - Alluxa ultranarrow filter 656.4nm, FWHM 0.1nm
- Collect light onto two 50μm fibers into ThorLabs SPADs
- Successful HBT measurement of fiber size
- Two 50μm fibers in one ferrule: “binary artificial star”
 - Rotate fiber bundle to measure fiber separation



$D_{\text{Measured}} = 51.5 \pm 1.0 \mu\text{m}$
 $D_{\text{Manufacturer}} = 50 \pm 1 \mu\text{m}$

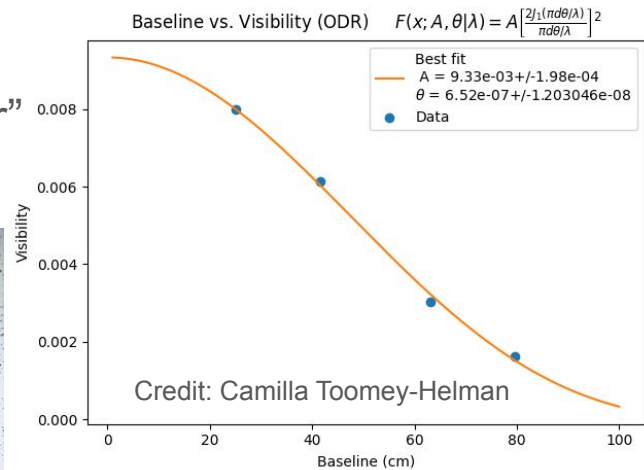
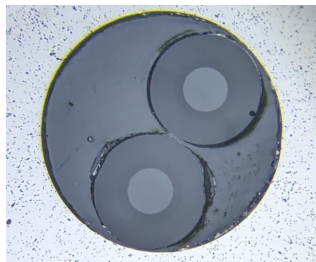
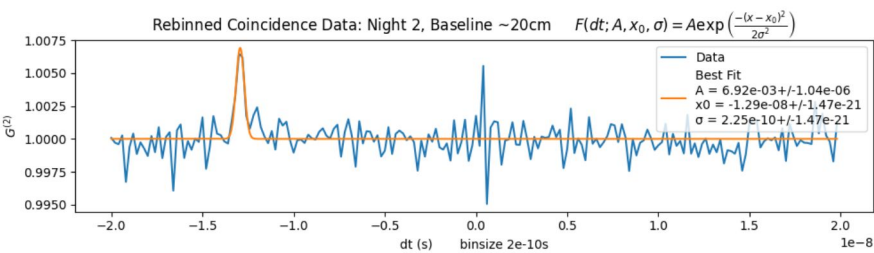
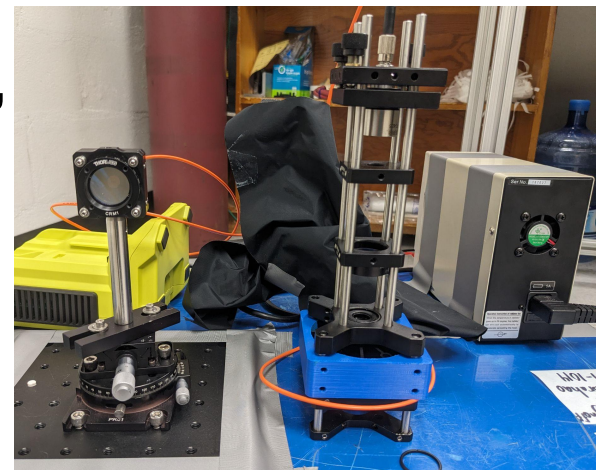
(<https://www.thorlabs.com/thorproduct.cfm?partnumber=M42L01>)

Baseline vs. Visibility (ODR) $F(x; A, \theta|\lambda) = A \left[\frac{2J_1(\pi x \theta / \lambda)}{\pi x \theta / \lambda} \right]^2$



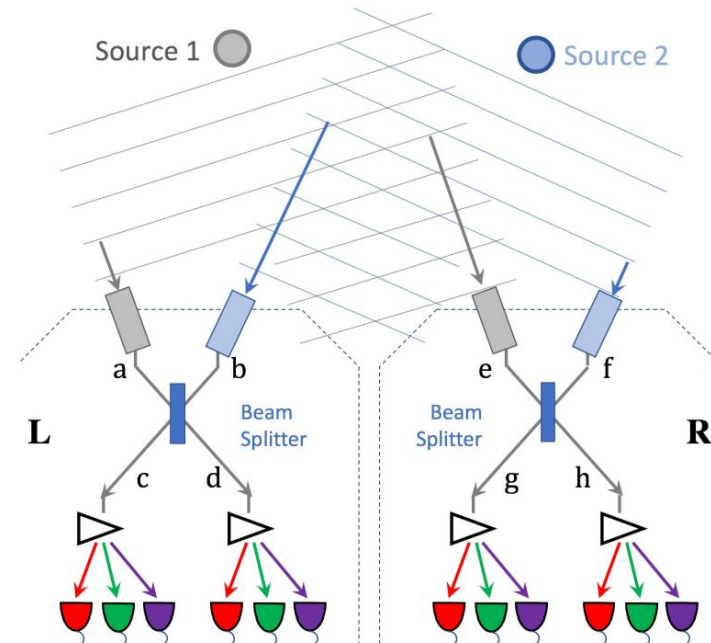
Ground Tests with Artificial LED “Star”

- “Artificial star”: $D=50\mu\text{m}$ fiber emitting filtered & polarized light from a red LED
 - Alluxa ultranarrow filter 656.4nm, FWHM 0.1nm
- Collect light onto two $50\mu\text{m}$ fibers into ThorLabs SPADs
- Successful HBT measurement of fiber size
- Two $50\mu\text{m}$ fibers in one ferrule: “binary artificial star”
 - Rotate fiber bundle to measure fiber separation



Conclusion

- Intensity Interferometer for wide-angle precision astrometry
 - Requires no optical connection between stations
 - Offline analysis
 - Improved baselines
 - Simplicity and scalability
- Active research:
 - Broadband HBT measurement
 - First-Order Adaptive optics
 - On the ground test measurements

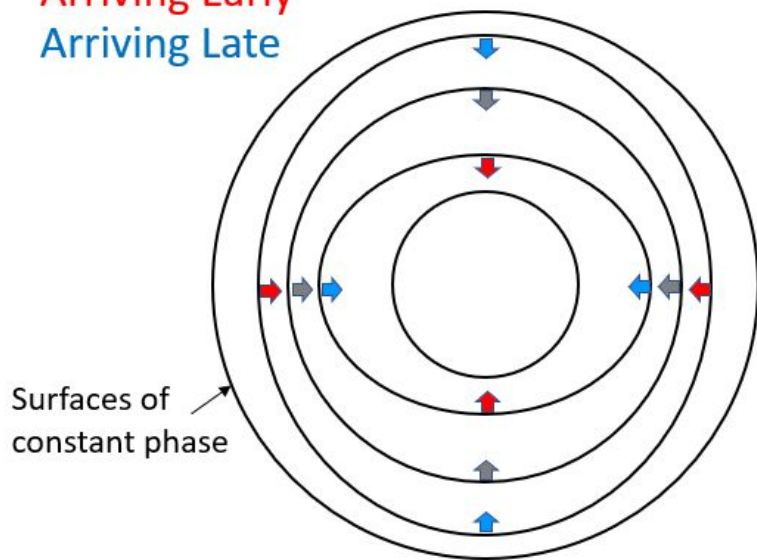


Stankus et al. 2022

<https://doi.org/10.21105/astro.2010.09100>

Friday 10:45 - Paul Stankus: Two-Photon Methods for Astrometric Gravitational Wave Detection

Arriving Early
Arriving Late



Pulsar Timing

Apparent Motion

The diagram shows five concentric circles representing surfaces of constant phase. A central grey arrow points outwards. Colored arrows (yellow, grey, green) are placed at various points on the circles, indicating the apparent motion of the wavefronts. A double-headed arrow labeled 'Apparent Motion' is shown above the circles, pointing towards the top-left.

Surfaces of constant phase



Astrometry

Acknowledgements

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Publications

J. Ferrantini, J Crawford, & S. Kulkov et al. 2024. arXiv:[2406.13959v1](https://arxiv.org/abs/2406.13959v1)

J. Crawford et al. 2023 *Optics Express* doi:[10.1364/oe.486342](https://doi.org/10.1364/oe.486342)

M. Keach et al. 2022 *Proc. SPIE 12183* doi:[10.1117/12.2632122](https://doi.org/10.1117/12.2632122)

P. Stankus et al. 2022 *Open Journal of Astrophysics* doi:[astro.2010.09100](https://doi.org/astro.2010.09100)

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