Piecing Together a Flat Hologram

Sabrina Pasterski

Celestial Holography in 30 minutes

The Celestial Conjecture:

scattering in asymptotically flat spacetimes is dual to a CFT living on the celestial sphere



This program evolved from a **bottom-up** approach to flat holography...



... recognizing **soft theorems as Ward Identities** for asymptotic symmetries and recasting the **soft operators as currents** in a codimension 2 CFT.

Soft Thm = Ward Id



 \Leftrightarrow

 $\langle out|Q^+[Y]\mathcal{S} - \mathcal{S}Q^-[Y]|in\rangle = 0$







In Bondi gauge the metric near future null infinity takes the form

$$ds^{2} = -du^{2} - 2dudr + 2r^{2}\gamma_{z\bar{z}}dzd\bar{z} + 2\frac{m_{B}}{r}du^{2} + (rC_{zz}dz^{2} + D^{z}C_{zz}dudz + \frac{1}{r}(\frac{4}{3}N_{z} - \frac{1}{4}\partial_{z}(C_{zz}C^{zz}))dudz + c.c.) + \dots$$
Radiative Data

which is preserved by the residual diffeomorphisms

$$\xi^{+} = (1 + \frac{u}{2r})Y^{+z}\partial_{z} - \frac{u}{2r}D^{\bar{z}}D_{z}Y^{+z}\partial_{\bar{z}} - \frac{1}{2}(u+r)D_{z}Y^{+z}\partial_{r} + \frac{u}{2}D_{z}Y^{+z}\partial_{u} + c.c$$

$$+ f^{+}\partial_{u} - \frac{1}{r}(D^{z}f^{+}\partial_{z} + D^{\bar{z}}f^{+}\partial_{\bar{z}}) + D^{z}D_{z}f^{+}\partial_{r}$$
Superrotations
Supertranslations

$$8\pi GQ^{+}[Y] = \int_{\mathcal{I}^{+}} \sqrt{\gamma} d^{2}z du \left[-\frac{1}{2} D_{z}^{3} Y^{z} u \partial_{u} C^{zz} + Y^{z} T_{uz} + u D_{z} Y^{z} T_{uu} + h.c. \right]$$
$$Q^{+}[Y] = Q_{S}^{+}[Y] + Q_{H}^{+}[Y]$$

$$\langle out|a_{-}(q)\mathcal{S}|in\rangle = \left(S^{(0)-} + S^{(1)-}\right)\langle out|\mathcal{S}|in\rangle + \mathcal{O}(\omega)$$
$$S^{(0)-} = \sum_{k} \frac{(p_k \cdot \epsilon^{-})^2}{p_k \cdot q} \qquad S^{(1)-} = -i\sum_{k} \frac{p_{k\mu}\epsilon^{-\mu\nu}q^{\lambda}J_{k\lambda\nu}}{p_k \cdot q}$$

Soft Thm = Ward Id



Soft Thm = Memory



Soft Thm = Memory



IR Triangle





Soft Theorems

4D Soft Mode = 2D Current

For a particular choice of Y

$$T_{zz} = 2iQ_S^+(Y^w = \frac{1}{z - w}, Y^{\bar{w}} = 0)$$

the superrotation Ward Id takes the form of a 2D stress tensor Ward Id.

$$\langle T_{zz}\mathcal{O}_1\cdots\mathcal{O}_n\rangle = \sum_{k=1}^n \left[\frac{h_k}{(z-z_k)^2} + \frac{\Gamma_{z_k z_k}^{z_k}}{z-z_k}h_k + \frac{1}{z-z_k}\left(\partial_{z_k} - |s_k|\Omega_{z_k}\right)\right]\langle \mathcal{O}_1\cdots\mathcal{O}_n\rangle$$



the asymptotic symmetry is physical



4D Soft Mode = 2D Current





4D Amplitude = 4D Correlator

LSZ \Leftrightarrow Extrapolate Dict.

$$\langle out|S|in\rangle_{boost} = \prod_{i} \lim_{r \to \infty} \int_{-\infty}^{\infty} \mathrm{d}\nu_{i} \,\nu_{i}^{-\Delta_{i}} \,\langle r\Phi(\nu_{1}, r, z_{1}, \bar{z}_{1})...r\Phi(\nu_{n}, r, z_{n}, \bar{z}_{n})\rangle$$

$$\nu = \{u, v\}$$



4D Amplitude = 2D Correlator



4D Amplitude = 2D Correlator

4D Lorentz invariance = 2D global conformal symmetry

$$\langle \mathfrak{G}_{\Delta_1}^{\pm}(z_1, \bar{z}_1) \dots \mathfrak{G}_{\Delta_n}^{\pm}(z_n, \bar{z}_n) \rangle = \prod_{i=1}^n \int_0^\infty d\omega_i \omega_i^{\Delta_i - 1} \langle out | \mathscr{S} | in \rangle$$

If we go to a boost basis, amplitudes transform as CFT correlators under the Lorentz group.

Collinear Limit = OPE

$$\begin{split} & \mathfrak{G}_{\Delta_1,+2}(z_1,\bar{z}_1)\mathfrak{G}_{\Delta_2,+2}(z_2,\bar{z}_2) \sim -\frac{\kappa}{2}\frac{\bar{z}_{12}}{z_{12}}B(\Delta_1-1,\Delta_2-1)\mathfrak{G}_{\Delta_1+\Delta_2,+2}(z_2,\bar{z}_2) + \dots , \\ & \mathfrak{G}_{\Delta_1,+2}(z_1,\bar{z}_1)\mathfrak{G}_{\Delta_2,-2}(z_2,\bar{z}_2) \sim -\frac{\kappa}{2}\frac{\bar{z}_{12}}{z_{12}}B(\Delta_1-1,\Delta_2+3)\mathfrak{G}_{\Delta_1+\Delta_2,-2}(z_2,\bar{z}_2) \\ & - \frac{\kappa}{2}\frac{z_{12}}{\bar{z}_{12}}B(\Delta_1+3,\Delta_2-1)\mathfrak{G}_{\Delta_1+\Delta_2,+2}(z_2,\bar{z}_2) + \dots , \end{split}$$



2D Radial Quantization → **More Symmetries**

For special weights, the SL(2,C) multiplets have primary descendants.

$$H^{k}(z, \bar{z}) := \lim_{\epsilon \to 0} \epsilon \, \mathfrak{G}_{k+\epsilon,2}(z, \bar{z}), \quad \Delta = k = 2, 1, 0, -1, \dots$$

Assuming these multiplets shorten, we have

$$H^{k}(z,\bar{z}) = \sum_{m=\frac{k-2}{2}}^{\frac{2-k}{2}} \bar{z}^{-\frac{k-2}{2}-m} H^{k}_{m}(z) , \qquad \qquad w^{p}_{n} = \frac{1}{\kappa} (p-n-1)! (p+n-1)! H^{-2p+4}_{n}(z) + \frac{1}{\kappa} (p-n-1)! (p+n-1)! (p+n-1)! H^{-2p+4}_{n}(z) + \frac{1}{\kappa} (p-n-1)! (p+n-1)! (p+n-1$$

2D Radial Quantization → More Symmetries

Complexifying the celestial sphere variables and defining a holomorphic commutator

$$[A,B](z) = \frac{1}{2\pi i} \oint_z dw A(w)B(z)$$

gives a $Lw_{1+\infty}$ symmetry algebra for appropriately rescaled modes

$$\left[w_{n}^{p}, w_{m}^{q}\right](z) = \left[n(q-1) - m(p-1)\right] w_{m+n}^{p+q-2}(z)$$

Celestial Algebra = Sym of SDG



Celestial Algebra = Chiral Algebra



Do these symmetries beyond tree level, or the self-dual sector?

Can we realize them in the matter sector?

Can we really complexify the celestial sphere to define these currents?



Celestial = Carrollian



perturbative bulk

Carrollian CFT₃

Celestial = Carrollian

 $ds^2 = -c^2 dt^2 + d\vec{x}^2$



Hard Charges = Light Ray Operators



Hard Charges = Light Ray Operators



Single Particle

Exclusive

VS

Inclusive









Kevin Costello Laurent Freidel Sabrina Pasterski Perimeter Institute for Theoretical Physics



Monica Pate NYU



Nima Arkani-Hamed IAS



Tim Adamo

University of Edinburgh



Lionel Mason Oxford



Natalie Paquette University of Washington



Andrew Strominger Jordan Cotler Harvard University



Tomasz Taylor Northeastern



Andrea Puhm École Polytechnique



David Skinner Cambridge

